This work was written as part of one of the author's official duties as an Employee of the United States Government and is therefore a work of the United States Government. In accordance with 17 U.S.C. 105, no copyright protection is available for such works under U.S. Law. Access to this work was provided by the University of Maryland, Baltimore County (UMBC) ScholarWorks@UMBC digital repository on the Maryland Shared Open Access (MD-SOAR) platform.

## Please provide feedback

Please support the ScholarWorks@UMBC repository by emailing <u>scholarworks-group@umbc.edu</u> and telling us what having access to this work means to you and why it's important to you. Thank you.

## Selecting highly-compact radio sources for the definition of the celestial reference frame

Patrick Charlot<sup>1</sup>, Alan L. Fey<sup>2</sup>, Roopesh Ojha<sup>2</sup>, David A. Boboltz<sup>2</sup>, J. I. B. Camargo<sup>3</sup> and A. Collioud<sup>1</sup>

<sup>1</sup>Observatoire de Bordeaux (OASU) - CNRS/UMR 5804, BP 89, F-33270 Floriac, France email: charlot@obs.u-bordeaux1.fr

> <sup>2</sup>Earth Orientation Department, U.S. Naval Observatory (USNO), 3450 Massachusetts Avenue NW, Washington, DC 20392-5420, USA email: afey,rojha,dboboltz@usno.navy.mil

<sup>3</sup>Observatório do Valongo, UFRJ, Lad. Pedro Antonio 43, Rio de Janeiro, RJ 20080-090, Brasil

Abstract. We discuss the issue of selecting highly-compact radio sources for the definition of the celestial reference frame.

**Keywords.** astrometry, reference systems, radio continuum: general, techniques: high angular resolution, techniques: interferometric, galaxies: active

The intrinsic radio structure of the extragalactic sources is one of the limiting errors in the definition of the International Celestial Reference Frame (ICRF). Based on multi-epoch VLBI images obtained with the Very Long Baseline Array and other VLBI telescopes around the world between 1994 and 2005, we evaluate this effect for 560 ICRF sources (about 80% of the current frame) and calculate a so-called 'structure index' to define the astrometric suitability of the sources.

The structure index ranges from 1 for the most compact sources to 4 for the most extended sources. The number of epochs for which the structure index is available for a given source varies from 1 for the least-observed sources to 20 for the intensively-observed sources. From this calculation, we identify a subset of 221 ICRF sources which have very good or good astrometric suitability (i.e., a structure index of either 1 or 2) at any of the available epochs.

We argue that these compact sources are potential candidates for defining the celestial frame with the highest accuracy when a future realization of the ICRF is made.