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Rapid Response Quality Control Service for the Laser Ranging Tracking Network

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Abstract Quality control systems for satellite laser ranging observations have been developed at a number of analysis institutes worldwide, using various software packages of precise orbit determination and data analysis. Satellite laser range observations, primarily from the two LAGEOS satellites but also from other satellites in low-Earth-orbits and up to GNSS altitude, are being processed on a sub-daily to weekly basis. The generated quality control reports are widely used to detect various kinds of problems and quickly provide anomalous information to laser ranging stations. They have been effective in shortening the time to return to normal when anomalous data are detected and in quantifying the performance of laser ranging stations. Consequently, a rapid feedback loop has now been incorporated in the modern satellite laser ranging operation.

Keywords Satellite laser ranging · Precise orbit determination · Quality control · Data anomaly detection

1 Introduction

The majority of currently active Satellite Laser Ranging (SLR) stations have achieved sub-centimeter precision in the two-way range measurements between a ground station and satellites at various altitudes. SLR has continuously contributed to global geodetic products, orbit validation and other areas, and, as a result, the high-quality range measurements in the optical regime have attracted many users.

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The quality control process, the topic of this paper, plays a key role in the operation of the SLR network. No systems are faultless especially for SLR because every SLR system is essentially different from others.

When an SLR station tracks a satellite and uploads the data, it cannot fully guarantee their quality. Stations have to calculate the root mean squares (RMS) of residuals generated during the normal-point formation process¹ (compressing all shot-by-shot data points to be represented by a single point per few seconds to a few minutes) so that they can reject obvious outliers. However, it is simply a statistical dispersion in its own observations, and generally speaking, it is not possible to locally assess the accuracy of their own observation data even at a meter or 10-meter level. Hence SLR stations need some feedback about the data quality based on independent analysis. This paper discusses the activities in the frame of the International Laser Ranging Service (ILRS) (Pearlman et al., 2002) to generate and disseminate quality control information to the tracking stations.

In the early stage of SLR, the observation data were recorded on magnetic tapes and it took usually a few weeks or even months to get them analyzed. When a problem occurred at a station, therefore, the anomalous data continued flowing in for a long time. By early 1990s, however, it became possible for the analysts to obtain the SLR data and provide feedback to the stations within a few days (Eanes et al. 1994; Ourensma and Noomen, 1998). This benefits from the progress of computer networking. Currently, ILRS tracking stations are encouraged to release the SLR data just after the end of the pass, usually within a few hours. We can thus apply a quick-look routine analysis within a day or even less.

There are a number of ongoing activities on SLR quality control in SLR analysis institutes, such as Glotov et al. (2004), Pavlis and Kuzmich-Cieslak (2007), Otsubo et al. (2008), Müller et al. (2013). In this paper, we shall focus on common procedures and worldwide collaboration, without going into the analysis details of each institute.

2 Data Flow

Continuous information flow in the SLR quality control procedure is outlined in Fig. 1. After the tracking data are transferred from each station to one of the two ILRS Operations Centers, data files are promptly organized and archived at the ILRS Data Centers after applying

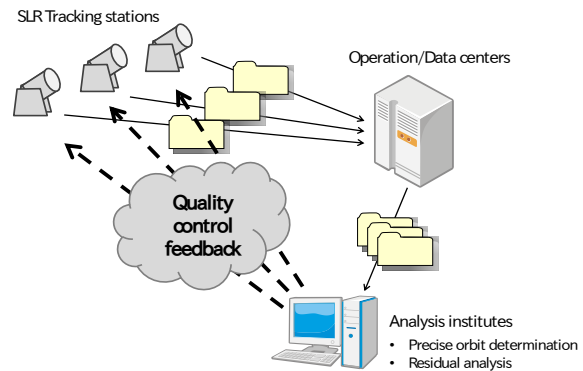


Fig. 1 SLR data flow in quality control sequence.

a format check and some nominal verification of the reported values in various fields (Noll et al. 2018).

The analysis institutes listed in Table 1 have routinely contributed to the quality control of the SLR data as of 2018. They routinely fetch the SLR data files from the Data Centers and pass them through a data reduction and orbit determination process to see how the observations fit the reference orbits. As shown in Fig. 1, the quality control institutes send quality control feedback to the tracking stations. In particular, when anomalous observations from a certain station are detected, the station should be notified.

Accuracy and reliability are the most important. Each institute has their own analysis software as listed in Table 1 and strive to improve its physical models and algorithms. Some of them have also benefited from long-term intercomparison tests through the Analysis Standing Committee of the ILRS (Luceri, 2018a). The analysis reports should be reliable and the analysis institutes make every efforts to avoid activating a false alarm.

Rapidness is also an important factor. Table 1 tells that the quality control analysis sequence is executed every day or every week. There are a couple of institutes that update the analysis reports even more frequently.

The results and products from these institutes are described in the next section.

3 Quality Control

3.1 Pass-by-pass analysis

Observational errors can be classified into two groups: random and systematic errors. The rapid quality control services for SLR are mainly focused on the detection of systematic error that far exceeds random error.

¹ https://ilrs.cddis.eosdis.nasa.gov/data_and_products/data/normal_point_system/

Table 1 Analysis institutes providing quality control information.

Institute	Software	Output	Satellites	Update	Duration
Astronomical Institute, University of Bern, Switzerland	Bernese 5.3	Range bias	GNSS	Daily	2000-present
Deutsches Geodätisches Forschungsinstitut, Germany	DOGS 5.4	Range & time bias	ETALON, LA-GEOS and LEOs	4-hourly	2003-present
Hitotsubashi University, Japan	c5++ R889	Range & time bias	GNSS, ETALON, LAGEOS and LEOs	6-hourly	1998-present
Joint Center for Earth Systems Technology, USA	GEODYN II and SOLVE II	Range & time bias, Residual map	ETALON, LA-GEOS and LEOs	Daily	2007-present
Information-Analytical Center, Russia	STARK-C 7.7	Range & time bias	LAGEOS	Daily	1997-present
NERC Space Geodesy Facility, UK	SATAN_SX	Residual plots	ETALON and LA-GEOS	Daily	1997-present
Shanghai Astronomical Observatory, China	SHORD-II	Range & time bias	ETALON and LA-GEOS	Weekly	1999-present
Wroclaw University of Environmental and Life Sciences, Poland	Bernese 5.2	Range bias & residual plots	GNSS	Daily	2016-present

Precise orbit determination is the main tool for the quality control. Every institute involved in quality control analysis has its own orbit determination software which is automatically run on a prescribed schedule. The majority of the institutes uses only SLR data to generate the reference (i.e. best-fit) orbits, but there are a couple of institutes (Astronomical Institute, University of Bern (AIUB), and Wroclaw University of Environmental and Life Sciences (WUELS)) whose reference orbits of the GNSS (Global Navigation Satellite System) satellites are generated from the microwave data.

For the purpose of quality control, the number of unknown parameters is limited. In the SLR-only analysis institutes, for instance, only orbital parameters are solved for while other parameters such as the coordinates of the stations are usually fixed to a-priori values. Therefore, the accuracy of the adopted station coordinates has a direct effect on the accuracy of quality control and it is very important, especially, in the case of new/re-installed stations, those significantly improved or those located in seismically active areas.

After iterative fits with outlier elimination, the post-fit residuals of normal points typically scatter from 0.7 to 1.5 cm weighted RMS for LAGEOS and LAGEOS-2. The residuals scatter more, from 1 to 4 cm weighted RMS, for low-orbit spherical satellites such as Starlette, Stella, Ajisai and LARES. The station-dependent weights are empirically based on the quality of the station data.

The top part of Fig. 2 shows typical post-fit residuals for LAGEOS 7-day SLR data, which is taken from the latest quality control analysis at Hitotsubashi University. 11 parameters (6 orbit (position and velocity)

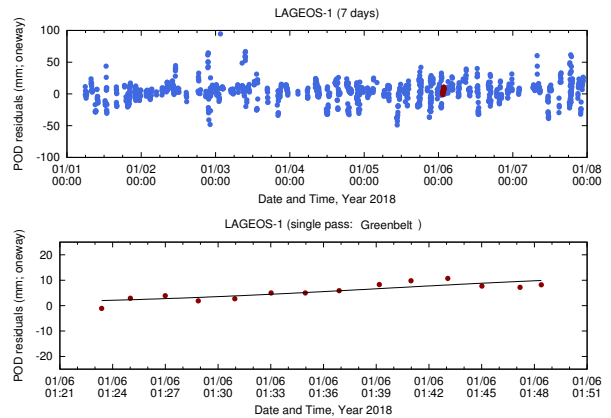


Fig. 2 A LAGEOS post-fit residual example of a good pass. Top: the whole network throughout the 7-day arc. Bottom: one of many good behaving passes.

elements and 5 empirical parameters in along-track and cross-track components) are solved for per 7-day arc. In this particular span, the weighted RMS was 1.0 cm after iterative removal of outliers. The remaining noise is stemmed from either imperfect physical models or observation errors. For instance, a pass observed from the Greenbelt SLR station is marked in red in both graphs of Fig. 2, and the bottom graph is simply zoomed-in with a 2-parameter best-fit curve (in black) of a range bias parameter and a time bias parameter, both of which are introduced later. Most of SLR data from good performing stations align near the zero level as in this example.

If we detect observations from a station deviating far from the zero level, i.e. the reference orbit, the station is likely to have a problem in its operation. We sometimes

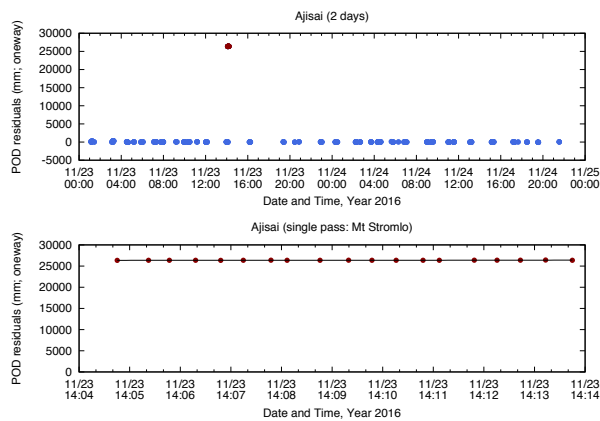


Fig. 3 An Ajsai post-fit residual example with large range bias. Top: the whole network throughout the 2-day arc. Bottom: a Mt Stromlo pass with +26 meter range bias.

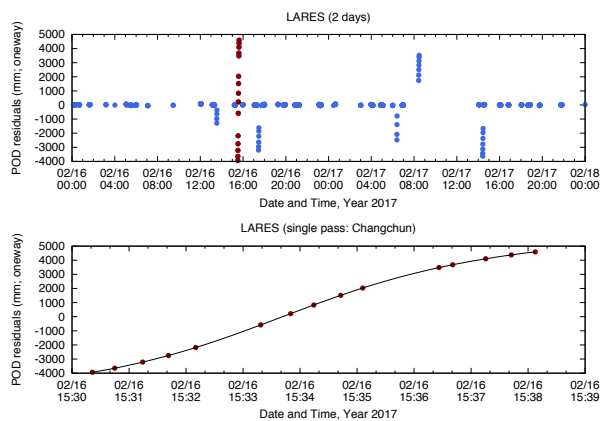


Fig. 4 A LARES post-fit residual example with large time bias. Top: the whole network throughout the 2-day arc. Bottom: a Changchun pass with +1 millisecond time bias.

see the cases like the example shown in Fig. 3 where all the SLR data in a pass are deviating greatly and consistently from the reference orbit. We also see the cases as the one displayed in Fig. 4 where the residuals show a large negative-to-positive (or opposite) trend.

These two kinds of errors are called range bias Δ_R and time bias Δ_T , respectively. The majority of anomalous data falls into either of them. The conceptual diagrams for a satellite pass are shown in Fig. 5. Range bias is defined as a constant error in the range observation and expressed in a unit of length such as meter or millimeter. Time bias is defined as a constant error in the time tag and expressed in a unit of time such as microsecond or millisecond. Note that the satellites dealt within this paper orbit the Earth at the velocity of roughly 4 to 8 km/s. A one-microsecond time bias thus corresponds to 4 to 8 mm in the along-track component.

Given a sufficient tracking duration and a sufficient number of normal-point observations, these two bias parameters, range bias Δ_R and time bias Δ_T , can be estimated from a population of post-fit residuals whose i -th element is defined as y_i . The observation equation is written as:

$$y_i = \Delta_R - \dot{\rho}_i \Delta_T + \epsilon_i \quad (1)$$

where $\dot{\rho}_i$ is the range rate for the i -th observation, and ϵ_i is the error in the i -th observation. This equation should be stacked for the number of observations per pass, and the two bias parameters Δ_R and Δ_T can be estimated in a standard least square adjustment.

In reality, it is sometimes difficult to estimate both of the two bias parameters when the tracking duration is too short or the number of observations is too few. In such cases, most analysis institutes provide only range bias, i. e. the simple mean of the residuals.

Fig. 6 is an example of quality control reports of Hitotsubashi University. The institutes in Table 1 whose output is ‘Range & time bias’ also generate and update numerical tables in similar formats. Each line corresponds to one pass. The column of ‘rb’ corresponds to estimated range bias, and that of ‘tb’ to estimated time bias, both of which are followed by their formal errors of Eq. 1 in parentheses. In this case of the Yarragadee station, the error in the orbit models, about 1 cm RMS, well exceeds the precision in the normal-point observations, about a few mm RMS, and therefore the bias values given here scatter more than the actual observation precision. This is the typical case for good observations from good stations. If these bias values showed a huge jump which is significantly larger than its estimated error, and if the bias persisted for multiple passes, it is likely that the station has a problem with their equipment, software or operation.

In addition to numerical tables, some institutes provide graphical information. Various kinds of plots are being generated in Deutsches Geodätisches Forschungsinstitut, Technische Universität München (DGFI-TUM). One example in Fig. 7 shows LAGEOS range bias estimates of the Herstmonceux station. Users can visually see long-term behavior of their daytime observations and the nighttime observations. In this particular case, due to the high latitude ($N 51^\circ$) of the station, the amount of day/night passes has a clear annual pattern, and the day-night differences of the moving average curve nicely stay within 3 mm. Graphical plots are clearly more comprehensible and more informative for users.

Historically, only the two LAGEOS satellites, i.e., LAGEOS and LAGEOS-2, had been applied to quality control since they are primarily used for geodetic anal-

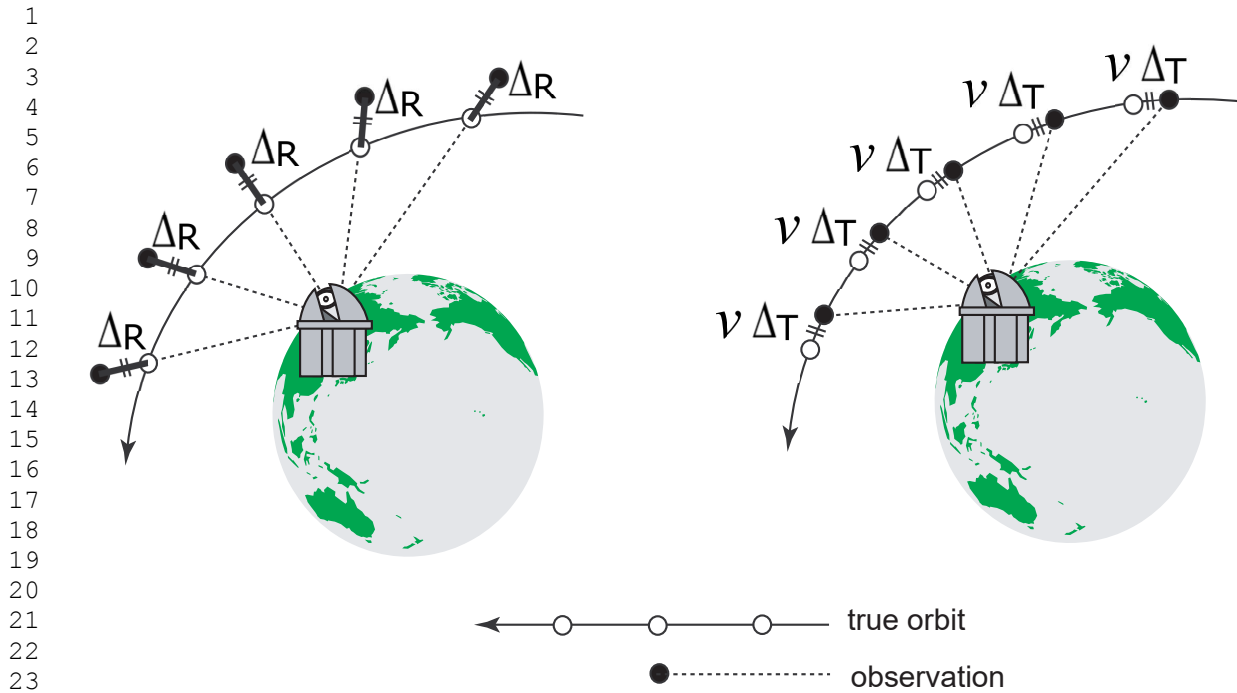


Fig. 5 Range bias (left) and time bias (right). Illustrated are positive biases in both cases. v is the velocity of the satellite.

```

# 7090 = YARRAGADEE
# sat site date time dur rb mm error tb us error prec bad total rms pres temp hum sdelay shft rms cfg r wlen
AJI1 7090 2018/01/07 18:54 7 14 ( 6 ) -4.9 ( 2.9 ) 3 0 / 17 9 979.5 295.8 31 13079 0 2 6 1 0 532
LARS 7090 2018/01/07 19:02 13 14 ( 2 ) -0.5 ( 0.5 ) 1 0 / 29 4 979.5 295.8 31 13079 0 2 6 1 0 532
GA03 7090 2018/01/07 20:48 132 -3 ( 6 ) 42.7 ( 24.3 ) 3 0 / 8 8 980.6 295.4 33 21896 -0 4 6 1 0 532
LAG1 7090 2018/01/07 21:01 21 6 ( 3 ) 2.3 ( 4.3 ) 3 0 / 10 5 980.2 294.9 32 13082 0 2 6 1 0 532
SARL 7090 2018/01/07 21:24 0 -11 ( 10 ) ----- - - - - - 2 0 / 4 3 980.3 294.6 33 13082 0 2 6 1 0 532
ETA1 7090 2018/01/07 21:48 57 1 ( 3 ) ----- - - - - - 6 0 / 11 14 980.8 295.2 35 21896 -0 4 6 1 0 532
SARL 7090 2018/01/07 23:02 5 8 ( 4 ) -3.3 ( 0.8 ) 1 0 / 13 4 981.0 296.4 33 13083 1 2 6 1 0 532
AJI1 7090 2018/01/07 23:09 2 20 ( 8 ) ----- - - - - - 3 0 / 7 11 981.0 296.4 33 13083 1 2 6 1 0 532
LAG2 7090 2018/01/07 23:15 19 17 ( 4 ) 1.4 ( 6.5 ) 2 0 / 11 4 981.0 296.9 33 13083 1 2 6 1 0 532
STRL 7090 2018/01/08 00:13 11 0 ( 3 ) -3.2 ( 0.7 ) 1 0 / 15 4 981.0 298.4 28 13084 0 2 6 1 0 532
LAG1 7090 2018/01/08 00:27 33 11 ( 2 ) 0.5 ( 1.5 ) 2 0 / 18 5 980.8 299.6 28 13084 0 2 6 1 0 532
AJI1 7090 2018/01/08 01:03 17 24 ( 5 ) 0.2 ( 1.0 ) 3 0 / 20 11 980.7 300.7 25 13084 0 2 6 1 0 532
CRY2 7090 2018/01/08 01:52 7 6 ( 4 ) 2.2 ( 0.7 ) 1 0 / 13 3 980.5 302.6 23 13084 0 2 6 1 0 532
STRL 7090 2018/01/08 02:04 7 9 ( 5 ) 6.9 ( 1.4 ) 2 0 / 11 5 980.4 303.1 22 13084 0 2 6 1 0 532
AJI1 7090 2018/01/08 03:12 1 18 ( 11 ) ----- - - - - - 2 0 / 4 9 979.9 305.0 20 13084 1 2 6 1 0 532
LAG2 7090 2018/01/08 03:17 32 1 ( 4 ) 5.7 ( 2.8 ) 3 0 / 12 5 979.7 305.7 19 13084 1 2 6 1 0 532
GA02 7090 2018/01/08 04:58 227 4 ( 3 ) -126.3 ( 54.0 ) 6 0 / 21 14 977.8 309.6 16 21897 1 6 6 1 0 532
GL33 7090 2018/01/08 06:42 103 74 ( 17 ) ----- - - - - - 7 0 / 4 17 976.8 310.5 15 21898 0 6 6 1 0 532
STEL 7090 2018/01/08 07:05 7 -1 ( 4 ) 0.8 ( 0.9 ) 2 0 / 17 6 977.1 310.8 16 13086 2 2 6 1 0 532
LAG1 7090 2018/01/08 07:19 9 10 ( 10 ) 9.8 ( 17.0 ) 4 0 / 6 4 976.9 310.7 16 13086 2 2 6 1 0 532
STRL 7090 2018/01/08 07:38 4 1 ( 4 ) 4.0 ( 1.9 ) 1 0 / 11 4 976.8 311.1 16 13086 2 2 6 1 0 532
LAG2 7090 2018/01/08 07:47 15 6 ( 8 ) 5.7 ( 8.5 ) 1 0 / 6 4 976.6 311.2 15 13086 2 2 6 1 0 532
LARS 7090 2018/01/08 07:55 3 -6 ( 8 ) 1.4 ( 2.3 ) 1 0 / 7 3 976.6 311.1 15 13086 2 2 6 1 0 532

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Fig. 6 Example of a quality control report (excerpted from the 16 January 2018 report issued by Hitotsubashi University). Listed are: satellite name, station ID, observation date and time, pass duration (in minutes), range bias and its estimation error (in mm), time bias and its estimation error (in microseconds), estimated precision of normal point (in mm), rejected and total data points, precision of raw ranging (in mm), pressure (in hPa), temperature (in Kelvin), humidity (in percents), applied system delay (in mm), delay shift (in mm), precision of calibration ranging (in mm), two ILRS configuration indicators, release flag and laser wavelength (in micrometers), from left to right.

yses, their characteristics are well understood and their orbit are less prone to poorly known physical process (e.g. atmosphere and earth gravity field). Analysis institutes have gradually added high-orbit and low-orbit satellites to the analysis of quality control reports. Despite the fact that the orbit determination precision of these satellites is not as good as that of the LAGEOS satellites, the use of a wide variety of satellites has great advantages as we explain in the following.

Firstly, the timing of problem occurrence can be exactly and reliably identified because of the increased number of passes. Secondly, we can detect problems that are not possible to be detected by analyzing only LAGEOS data because some stations have multiple hardware configurations for different satellite altitudes and because some stations are more productive in lower or higher satellites than LAGEOS. Thirdly, a range-dependent bias, often proportional to the range resulting from a

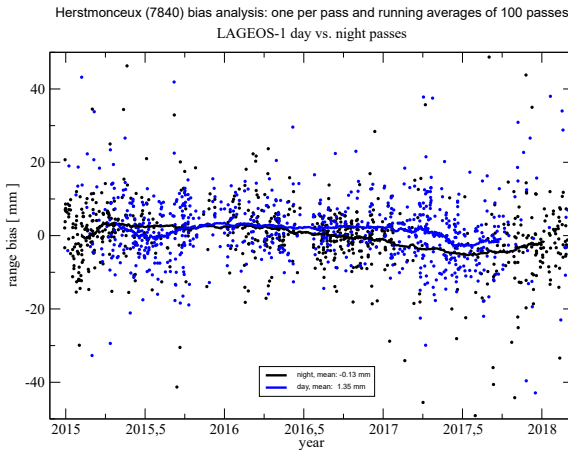


Fig. 7 A plot example of LAGEOS range bias of the Herstmonceux station, UK. DGFI-TUM routinely generates not only numerical tables but also these plots for all stations.

problem in the clock frequency, can be clearly identified.

Sufficient coverage during a pass is critical to this kind of quality control. With just a few normal-point observations over a short time, it is not possible to estimate reliably the two parameters. As a consequence, we often estimate the range bias only. It should be noted that long duration SLR passes are important for data quality assessment. In particular, a horizon-to-horizon pass duration of a high-orbit target, such as the two ETALON satellites and the GNSS satellites, is typically 5 to 8 hours. We need SLR coverage of at least 1 hour to separate the two bias parameters. However, when the SLR coverage during a pass is short, we often find that the time bias parameter cannot be determined at all, or are not well determined, for high-orbit satellite passes.

It is no surprise therefore that AIUB provides only the range bias estimates in their routine reports dedicated to the high-orbit GNSS satellites. Reference orbits in the AIUB quality control analysis are based only on GNSS microwave data. In 2017, the number of registered passes to GNSS was almost four times larger than the number of LAGEOS passes, which allows for a rigorous analysis of SLR biases at high-altitude targets. The AIUB orbits come from the analysis at the International GNSS Service processing at the Center for Orbit Determination in Europe (Dach et al. 2009). Consequently, the reported SLR bias values include the systematic offset between the two independent techniques: Thaller et al. (2011) studied the impact of mismodelled antenna phase center issues and unconsidered range bias, and Sošnica et al. (2015) investigated the patterns of SLR

residuals in relation to SLR system types and onboard retroreflector array types.

As a result of there being a lot more SLR satellites in space, a ground station can now see 10 SLR satellites or even more above the horizon while it can track only one satellite at a time. The stations are encouraged to switch targets frequently (often called “interleaving mode”) rather than solid tracking from the beginning to the end of the pass, and, by doing so, they would obtain SLR data from more satellites in an efficient way.

3.2 Residual plots

Quality control information is reduced to just one line per pass in the previous subsection. We have shown several normal-point residual graphs (Figs. 2, 3 and 3), but they are not routine products. Residual graphs are indeed more informative and useful for a precise understanding of the behavior of observation data. NERC Space Geodesy Facility (SGF) and WUELS provide online residual graphs so that the users can visually see the precise details of normal-point residuals such as outliers and trends.

The SGF updates on its website each day interactive normal-point residual plots of LAGEOS and Etalon normal points submitted globally over the past 7 days as a by-product of its routine analysis center activity. The SGF also separately carries out daily “short-arc” quality control, which extracts and plots a common-view pass from multiple stations located nearby (Sinclair and Appleby, 1993). Through these plots available online, we see how each normal point behaves with respect to the reference orbit. An example is shown in Fig. 8, which shows four stations tracking LAGEOS. In this particular case, the LAGEOS satellite flies firstly over Europe and then over the USA. This plot would suggest a range bias is present in the Borowiec range measurements. These plots show solid/sparse tracking patterns and the behavior of each normal point. For instance, in the past, we frequently saw outliers at the beginning or at the end of a pass, but it is rare today thanks to alerts from this kind of service.

WUELS provides online tools for an interactive generation of SLR residual plots of GNSS satellites, based on the microwave orbits processed at AIUB in the framework of the IGS multi-GNSS Experiment (Prange et al. 2017). More than 50 active GNSS satellites are today tracked by the SLR stations, which results in a substantial number of SLR observations. The WUELS service allows for the residual analysis with respect to the date and time of data acquisition, differences from nighttime and day-time tracking, dependencies of residuals as a function of the elevation and azimuth angles. All

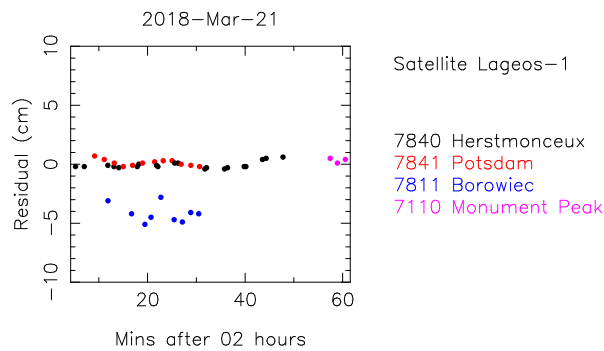


Fig. 8 A residual plot example of a LAGEOS-1 common-view pass from Europe-USA region. This is taken from NERC SGF's daily short-arc quality control.

of them help in identification of various systematics. The analyses can be performed for individual stations and satellites or for a user-defined group of satellites or stations.

4 Integration and communications

4.1 Quality control feedback

All of the quality control activities described in the previous section are intended to give prompt feedback to the tracking stations. Some stations are routinely watching the quality control reports, and the analysis institutes also send an alert message when they detect anomalous observations.

All of the analysis institutes in Table 1 routinely update the quality control reports on their websites and/or make them available through the ILRS archive and the ILRS mailing list. The long-time records of daily and weekly reports are archived not just at each institute but also aggregated on the ILRS Data Center, at Crustal Dynamics Data Information System (CD-DIS) of National Aeronautics and Space Administration (NASA). The URLs of the SLR quality control services are collectively listed in Table 2.

Users can see the quality control results from multiple institutes at once. Range bias estimates of the two LAGEOS satellites from multiple institutes are easily compared by the combined range bias reports weekly sent by AIUB. Each line of the combined report contains range bias estimates from five analysis institutes. Laser station operators can verify the reliability by checking whether an anomalous result is detected by all institutes or just by one or few institutes. It also helps analysts to see whether their results are in harmony with others.

It has been common to look at the text-based reports in the analysis institutes and also in the tracking

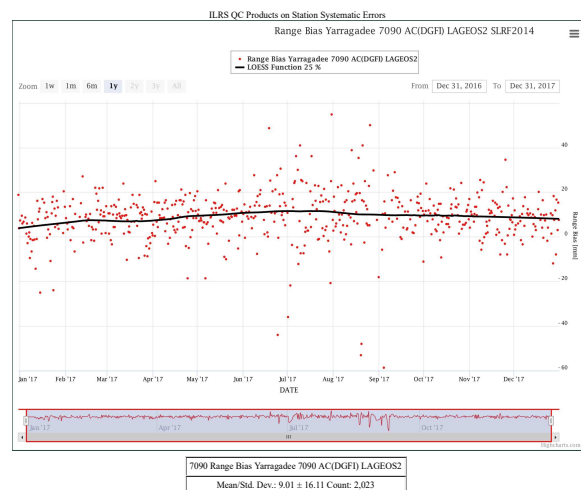


Fig. 9 A range bias plot example of LAGEOS-2 range bias estimates of Yarragadee in 2017, plotted by JCET's online interactive tool. DGFI-TUM's quality control results are chosen in this case.

stations, but an interactive graphical web tool helps more. The time series of pass-by-pass range bias, time bias and other quality control parameters can flexibly be plotted in a web browser using the QC-viewer, a service provided by the Joint Center for Earth Systems Technology (JCET) as shown in Fig. 9. This service helps users to visually understand the long-term and short-term behaviors of the quality of a certain station, optionally with a smoothing curve. The user has various options to choose from, such as the satellite, the station, the quantity of interest, the axis scales and so on, for customized graphs. In addition to JCET's own analysis results users can also choose those from the other QC institutes appearing on Table 1.

The analysis institutes look through the quality control reports. When a series of anomalous observations are clearly detected, it is very likely stem from the observation itself, not from the analysis side. A warning message is then sent to the station. The whole quality control procedure in the analysis and the data transfer is almost automated, but manual handling is still to some extent involved in detecting anomalous data and also notifying the stations.

We usually use email for communication with stations. In particular, DGFI-TUM, Hitotsubashi University and JCET have been active so far in sending alerts to the worldwide stations. In June 2011, the analysis institutes agreed to use the newly established "Rapid-ServiceMail" mailing list implemented at DGFI-TUM, which is useful to notify and archive the anomaly detection events among analysis institutes.

In addition to the worldwide services, there are also regional activities for SLR quality control. For instance,

Table 2 URLs for the quality control services.

Institute/Service	URL
Individual	
DGFI-TUM	https://ilrs.dgfi.tum.de/quality/weekly_biases/stations/
Hitotsubashi University	http://geo.science.hit-u.ac.jp/slr/bias/
NERC SGF	http://sgf.rgo.ac.uk/
WUELS	http://www.govus.pl/
Integrated	
ILRS Data Center, CD-DIS, NASA	ftp://cddis.gsfc.nasa.gov/pub/reports/
ILRS Global Performance Cards	https://ilrs.cddis.eosdis.nasa.gov/network/system_performance/global_report_cards/monthly/
ILRS RapidService-Mail, DGFI-TUM	http://rapidservicemail.dgfi.tum.de/
JCET QC-viewer	http://geodesy.jcet.umbc.edu/QC/
JCET Visualization of ILRS Report Cards	http://geodesy.jcet.umbc.edu/ILRS_REPORT_CARD/

the Information-Analytical Center of Russia provides weekly feedback to stations of the Russian SLR network in the Russian language. Their weekly reports sent to the Russian SLR stations contain quantitative and qualitative statistical information customized for the Russian network, which includes the widely-available pass-by-pass quality control (range bias and time bias) reports of the LAGEOS satellites.

It is now possible to detect and notify a problem within a day. If a problem can be subsequently corrected afterwards, the station is strongly advised to resubmit the corrected observations with an incremented release flag. We have seen that most of the problems can be solved instantly or within a few days. It minimizes the time that large bias is present in the data.

Pass-by-pass analysis reports from the analysis institutes are gathered at the ILRS Central Bureau where a long-term statistics report is published from both qualitative and quantitative aspects. The ‘‘SLR Global Performance Report Card’’ covers the latest 12 months and is currently updated every month (every 3 months until 2012). It consists of two tables: one is for quantity and the other is for quality. The first table provides the ranking by the number of passes, and the number of normal point observations, for different kinds of satellites. The second table provides a summary of the five quality control reports providing a pass-by-pass bias analysis of the LAGEOS satellites. The results are averaged over time to produce a stability measure in two ways: a pass-by-pass range bias variation over the last three months (‘short-term’ stability) and a monthly-average range bias variation using over the last twelve months (‘long-term’ stability). These tables are helpful to compare a station’s performance with respect to others and often encourage the improvement of the quality and quantity of their data. They are also useful to mon-

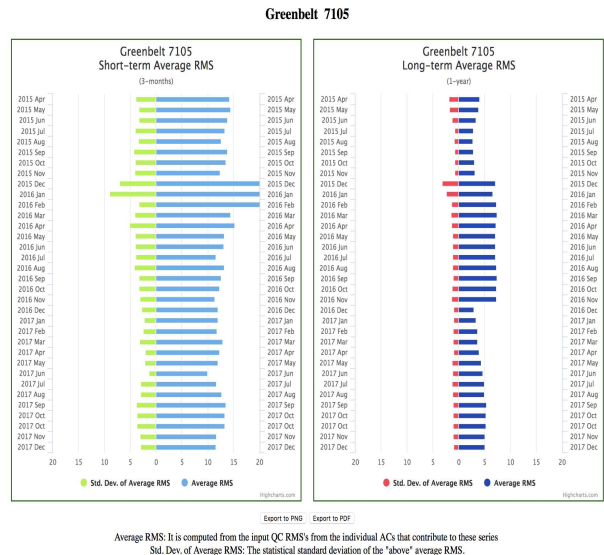


Fig. 10 Rendering of the short-term and long-term stability measure from the quarterly report card of Greenbelt using the online interactive tool developed by JCET.

itor the evolution of a station’s performance over time. To facilitate these comparisons, JCET has established a service that archives all report cards’ data online and allows the user to create plots of these statistics over time for each station of interest (Fig. 10). The results plotted are the average of the available short-term or long-term stability estimates from all quality control centers and the standard deviation of that estimate.

4.2 Causes of anomalous measurements

We often receive a reply from the station about the cause of the problem. Based on our long experience in

1 interacting with the stations, we have learned that cer-
2 tain patterns are connected to specific causes.

3
4 The most frequent one is constant offset range bias.
5 Some SLR laser systems transmit multiple pulses at a
6 fixed interval within one or a few nanoseconds, where
7 one pulse is usually predominantly stronger than oth-
8 ers. However, especially in the case that there is more
9 than one strong pulse, the choice of the main pulse
10 could be wrong. The other major cause originates from
11 calibration. It is often caused by reflection from a point
12 different from the calibration target. If the internal sys-
13 tem delay is wrongly measured, the SLR data are shifted
14 by the same amount as its error. In both cases, a jump
15 of the system delay is also observed. It is thus impor-
16 tant to always keep an eye on the time series of the
17 system delay measurement. Human input error during
18 operation or data reduction can also be the cause.

19
20 Time bias is also detected almost as often as range
21 bias. Note that the SLR observables are basically the
22 combination of a time tag and a two-way range. Large
23 time bias indicates a problem in time tagging. Time
24 bias sometimes appears constant, and sometimes vary-
25 ing, i.e. the pass-by-pass time bias estimates gradually
26 change over time. The problem comes from a frequency
27 source, other timing devices and sometimes software.
28 In the past, one second time bias was often seen after
29 a leap second but it is rare nowadays. Extremely large
30 time bias of one day (wrong day) and even one year
31 (wrong year) can be also detected at times.

32
33 Meteorological sensors such as a barometer and a
34 thermometer are sometimes found to be an error source.
35 Atmospheric delay models are constructed as a function
36 of meteorological data and therefore a problem in these
37 instruments manifests itself as a range error.

38
39 Analysis institutes must pay attention not to send a
40 false alarm. Orbit determination has not always prop-
41 erly converged due to an insufficient data amount, too
42 many outliers, imperfect physical models, and orbit ma-
43 neuvers. For instance, assume that a satellite is sparsely
44 tracked even by the whole network and one productive
45 station has large time bias, and then a well-performing
46 station located nearby sometimes has spurious time bias
47 estimates too. This is because time bias is insepara-
48 ble from the along-track component of satellite orbits,
49 and because large time bias can affect the orbital pa-
50 rameters if the actual anomalous data cannot be prop-
51 erly eliminated in the orbit determination stage. In an-
52 other instance, the use of accurate station positions
53 and velocities is essential for reliable and precise qual-
54 ity control results. Analysis institutes thus obtain, or
55 update, the station coordinates if a station is new, has
56 undergone significant changes, or is displaced due to
57 earthquakes. Currently, the latest version of SLRF2014

(Luceri et al. 2018b), an extension of ITRF2014, is
widely adopted for quality control analyses and fre-
quently updated.

5 Conclusions and ongoing/future studies

Precise orbit determination techniques and high-speed
communications have been well integrated to realize
rapid quality control feedback for the SLR station net-
work. Analysis institutes have routinely provided anal-
ysis reports in various ways and they have contributed
to improving the overall quality and stability of the SLR
observations.

A new ILRS study group called Quality Control
Board was newly organized in December 2015. The
rapid quality control activities covered in this paper
are one of the core tasks in this group, and the latest
news and future plans are being discussed in monthly
international teleconferences.

High-quality SLR data is beneficial not only for geodesy
but also for satellite altimetry, satellite navigation and
any mission that relies on SLR for its precise orbits.

This paper focused on the rapid response feedback
for relatively large systematic errors, but it is also pos-
sible to look into systematic observation errors/trends
at the millimeter level by accumulating longer span of
SLR data (Otsubo et al. 2014; Appleby et al. 2016).
Both quality control methods, i.e. rapid and precise,
should be utilized in a complementary fashion in the
future so that systematic errors of any magnitude are
eliminated from the data.

A new technology, Time Transfer by Laser Link
(T2L2) demonstrated on the satellite Jason-2, makes
it possible to examine the accuracy of time tagging at
SLR stations (Exertier et al. 2017) with a 1 ns accuracy.
This is much higher than the requirement for ordinary
SLR, roughly 100 ns for 1 mm ranging accuracy, and
several anomalous cases have been detected. The T2L2
mission on Jason-2 was terminated in April 2018, and
the future missions are strongly desired from the view-
point of quality control.

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Fig. 1

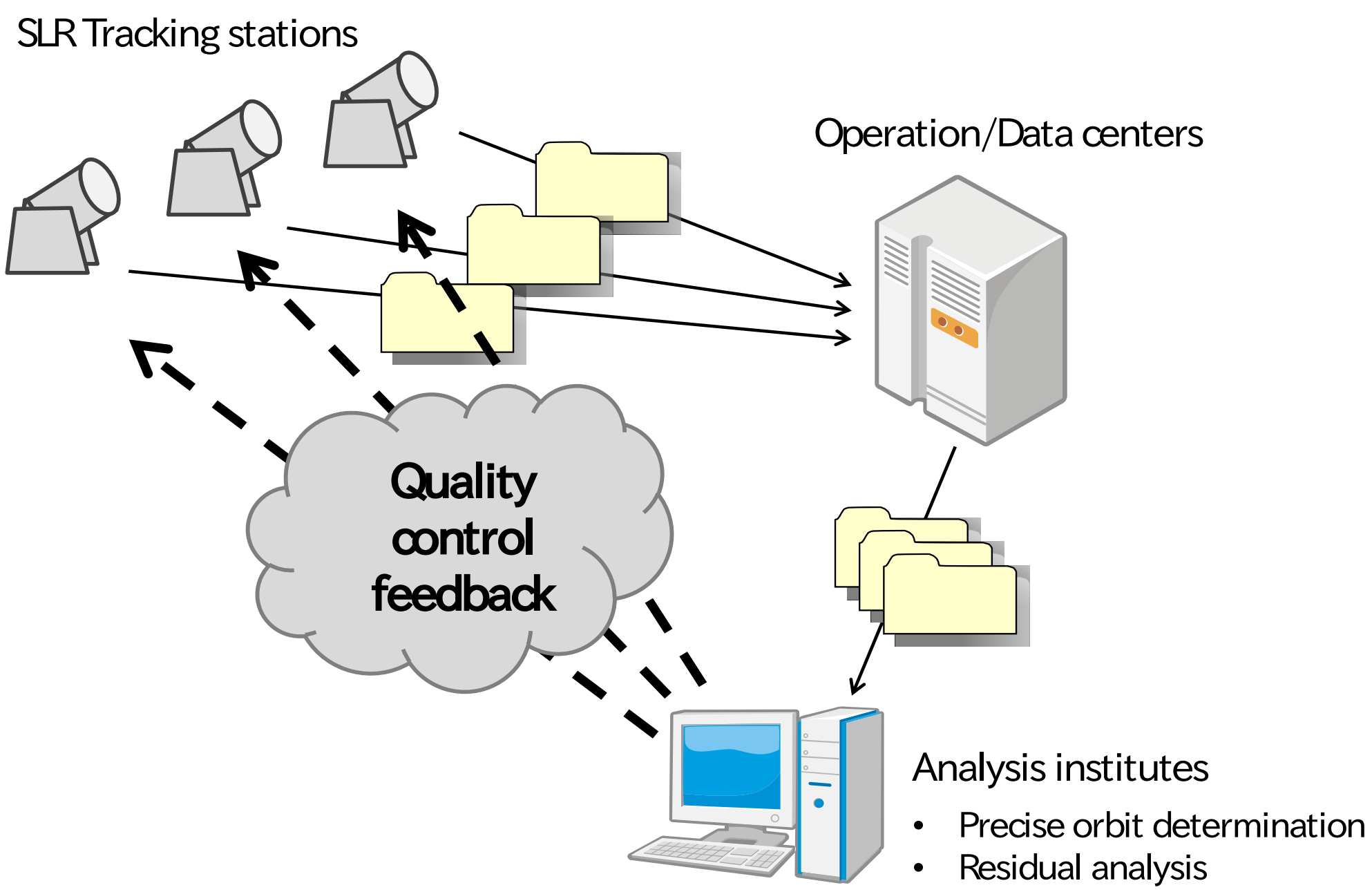


Fig. 2

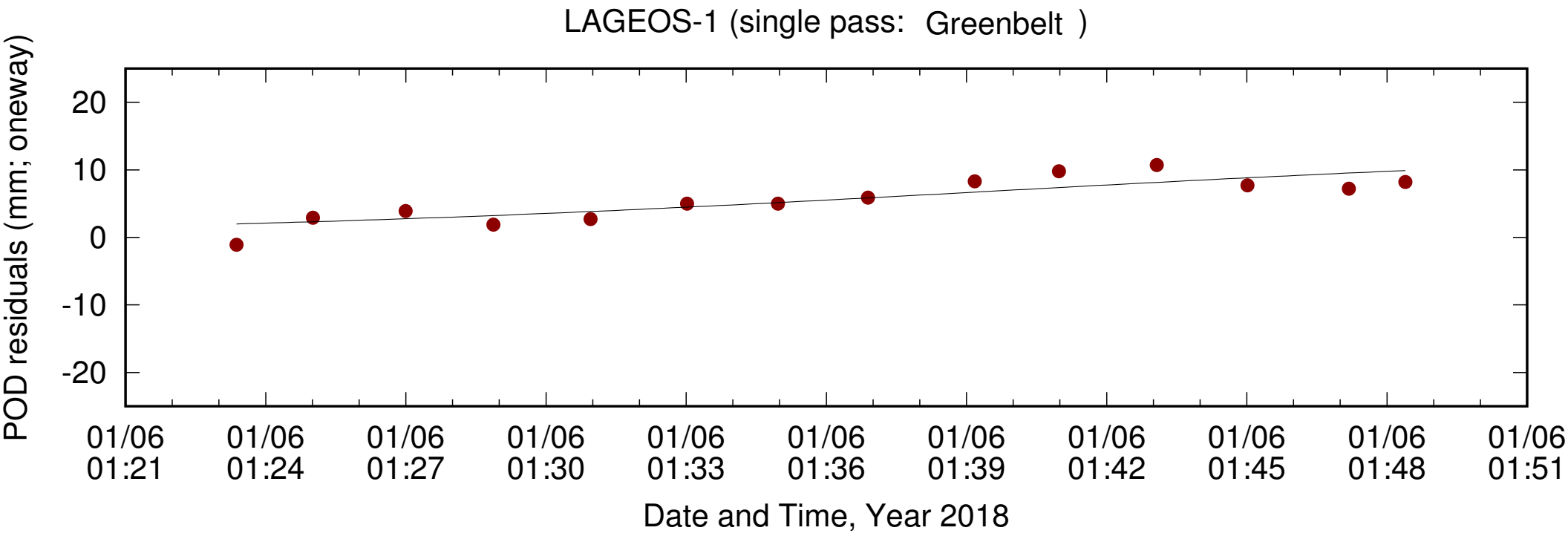
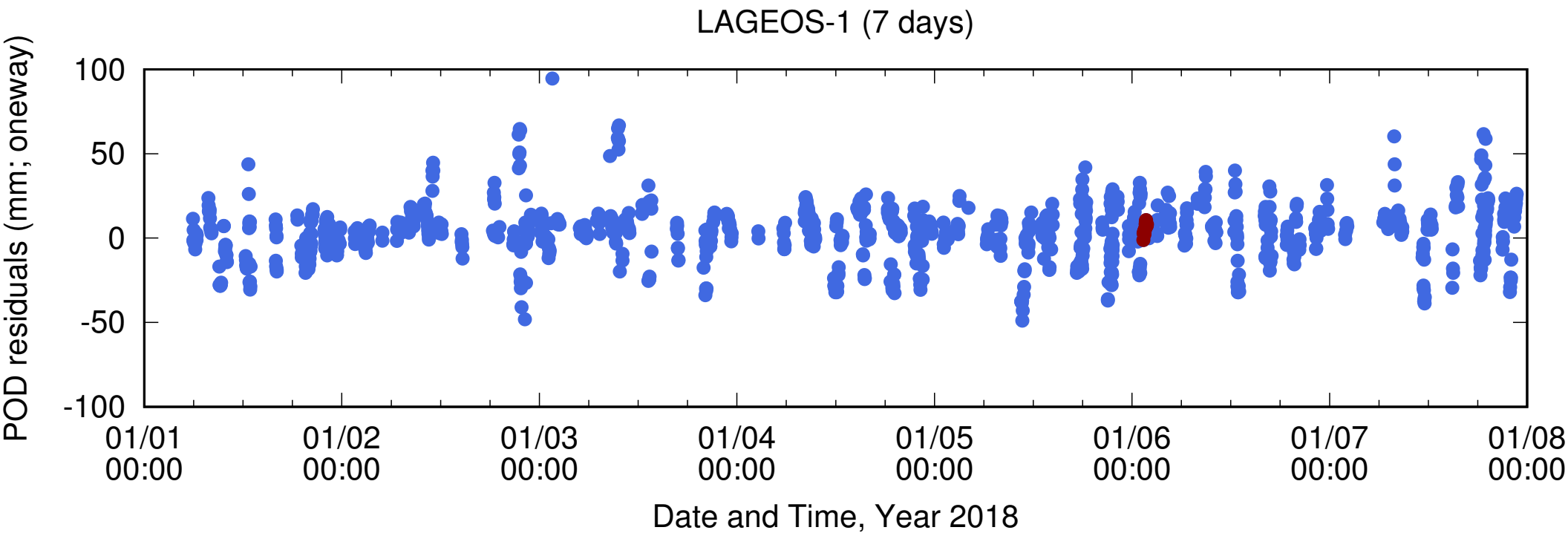


Fig. 3

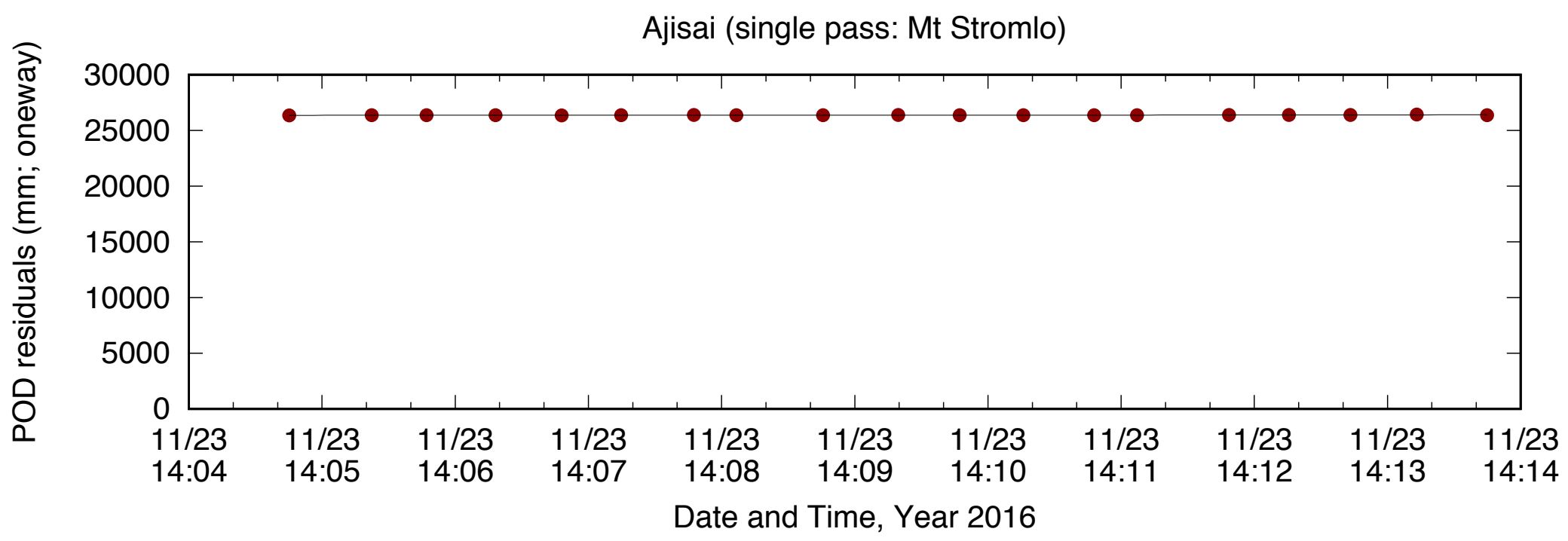
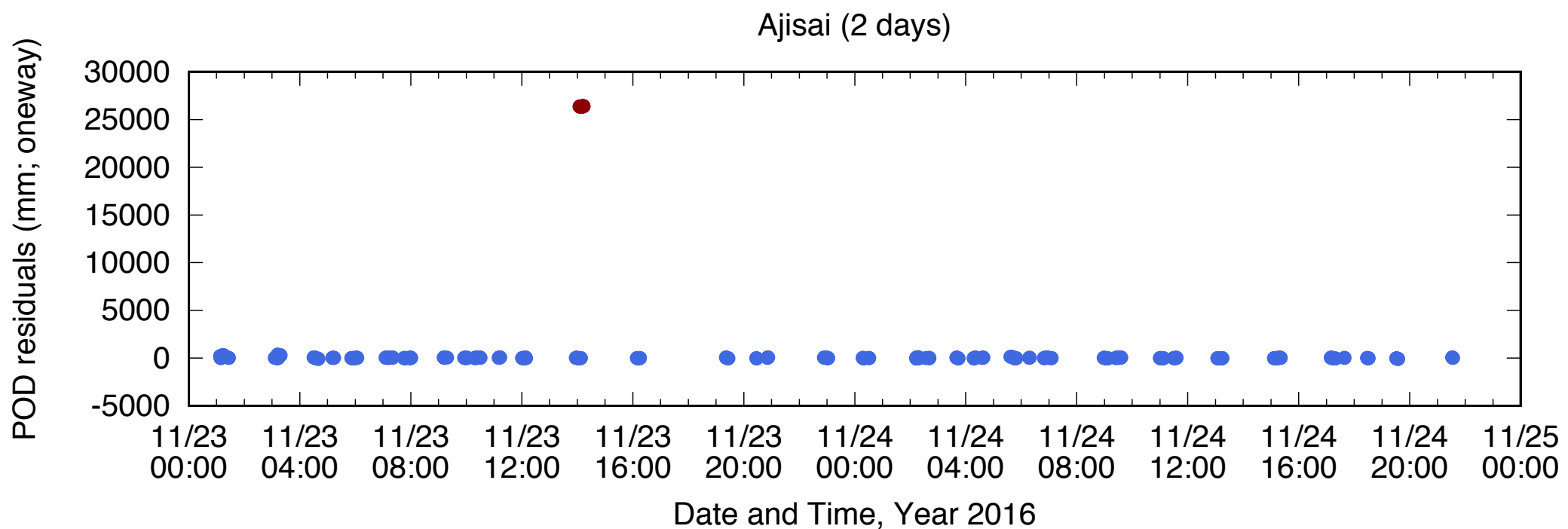


Fig. 4

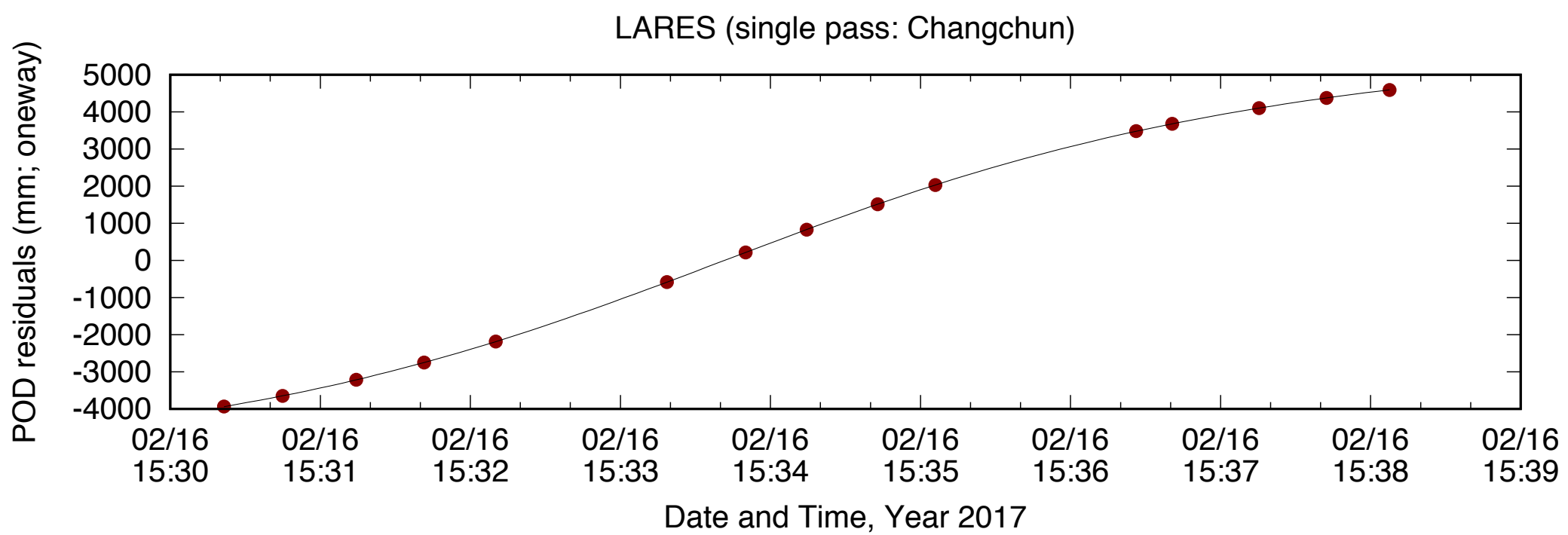
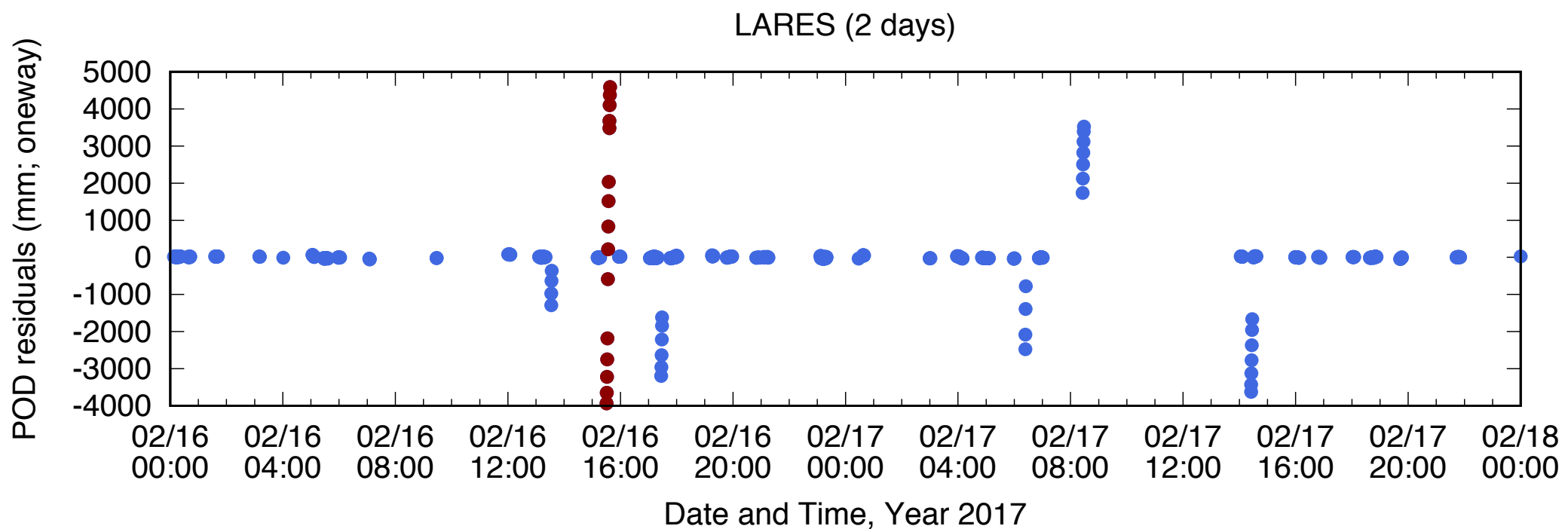


Fig. 5

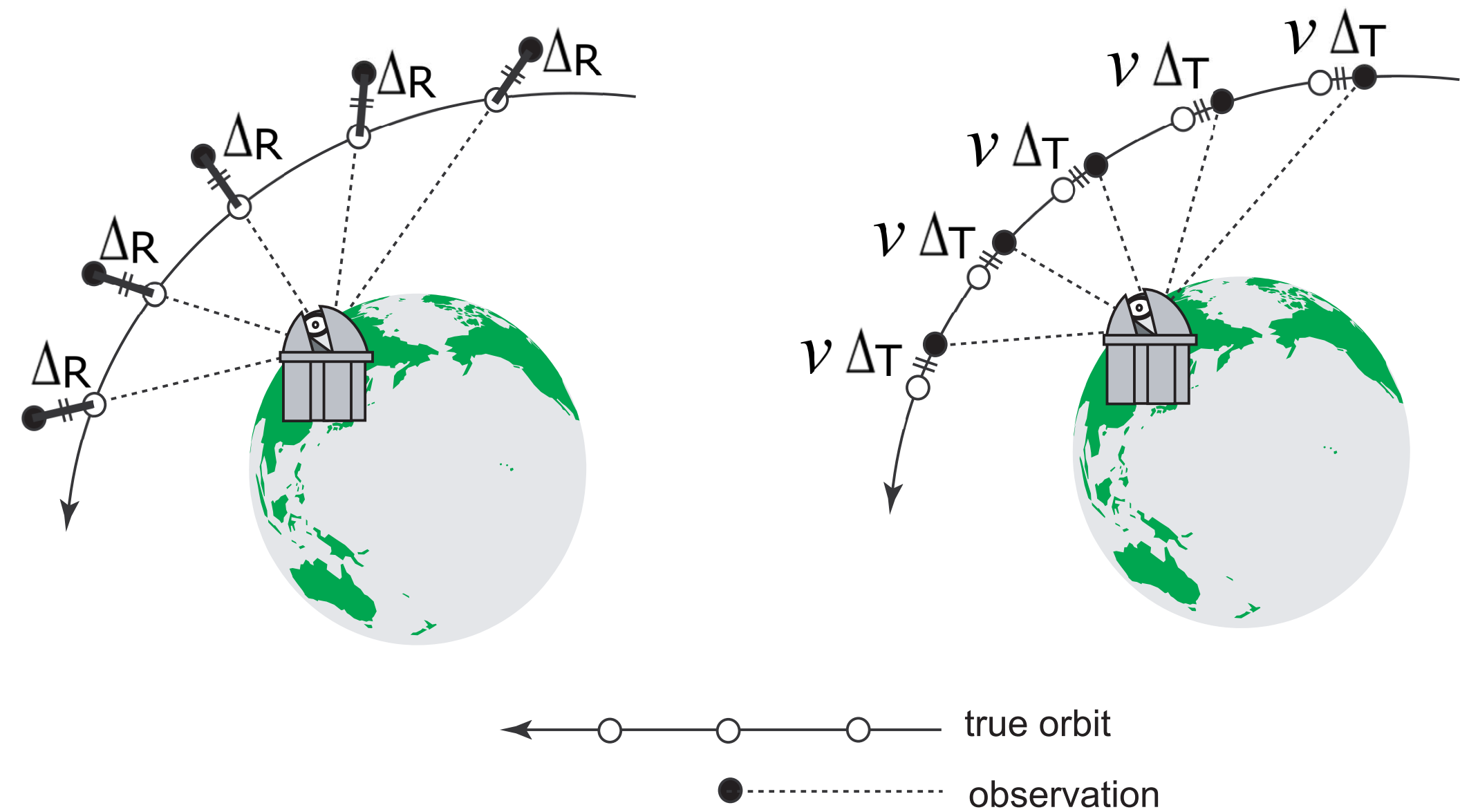


Fig. 6

#	#	#	sat	site	date	time	dur	rb	mm	error	tb	us	error	prec	bad	total	rms	pres	temp	hum	sdelay	shft	rms	cfg	r	wlen	
				7090 = YARRAGADEE																							
			AJ11	7090	2018/01/07	18:54	7	14		(-4.9	(2.9)			9	979.5	295.8	31	13079	0	2	6	1	0	532
			LARS	7090	2018/01/07	19:02	13	4		(-0.5	(0.5)			4	979.5	295.8	31	13079	0	2	6	1	0	532
			GA03	7090	2018/01/07	20:48	132	-3		(42.7	(24.3)			8	980.6	295.4	33	21896	-0	4	6	1	0	532
			LAG1	7090	2018/01/07	21:01	21	6		(2.3	(4.3)			5	980.2	294.9	32	13082	0	2	6	1	0	532
			SARL	7090	2018/01/07	21:24	0	-11		(---	(---)			3	980.3	294.6	33	13082	0	2	6	1	0	532
			ETAI	7090	2018/01/07	21:48	57	1		(---	(---)			14	980.8	295.2	35	21896	-0	4	6	1	0	532
			SARL	7090	2018/01/07	23:02	5	8		(-3.3	(0.8)			14	981.0	296.4	33	13083	1	2	6	1	0	532
			AJ11	7090	2018/01/07	23:09	2	20		(---	(---)			11	981.0	296.4	33	13083	1	2	6	1	0	532
			LAG2	7090	2018/01/07	23:15	19	17		(1.4	(6.5)			4	981.0	296.9	33	13083	1	2	6	1	0	532
			STRL	7090	2018/01/08	00:13	11	0		(-3.2	(0.7)			4	981.0	298.4	28	13084	0	2	6	1	0	532
			LAG1	7090	2018/01/08	00:27	33	11		(0.5	(1.5)			5	980.8	299.6	28	13084	0	2	6	1	0	532
			AJ11	7090	2018/01/08	01:03	17	24		(0.2	(1.0)			11	980.7	300.7	25	13084	0	2	6	1	0	532
			CRY2	7090	2018/01/08	01:52	7	6		(2.2	(0.7)			3	980.5	302.6	23	13084	0	2	6	1	0	532
			STRL	7090	2018/01/08	02:04	7	9		(6.9	(1.4)			5	980.4	303.1	22	13084	0	2	6	1	0	532
			AJ11	7090	2018/01/08	03:12	1	18		(---	(---)			9	979.9	305.0	20	13084	1	2	6	1	0	532
			LAG2	7090	2018/01/08	03:17	32	1		(5.7	(2.8)			5	979.7	305.7	19	13084	1	2	6	1	0	532
			GA02	7090	2018/01/08	04:58	227	4		(-126.3	(54.0)			14	977.8	309.6	16	21897	1	6	6	1	0	532
			GL33	7090	2018/01/08	06:42	103	74		(---	(---)			17	976.8	310.5	15	21898	0	6	6	1	0	532
			STEL	7090	2018/01/08	07:05	7	-1		(0.8	(0.9)			6	977.1	310.8	16	13086	2	2	6	1	0	532
			LAG1	7090	2018/01/08	07:19	9	10		(9.8	(17.0)			4	976.9	310.7	16	13086	2	2	6	1	0	532
			STRL	7090	2018/01/08	07:38	4	1		(4.0	(1.9)			4	976.8	311.1	16	13086	2	2	6	1	0	532
			LAG2	7090	2018/01/08	07:47	15	6		(5.7	(8.5)			4	976.6	311.2	15	13086	2	2	6	1	0	532
			LARS	7090	2018/01/08	07:55	3	-6		(1.4	(2.3)			3	976.6	311.1	15	13086	2	2	6	1	0	532

Herstmonceux (7840) bias analysis: one per pass and running averages of 100 passes
LAGEOS-1 day vs. night passes

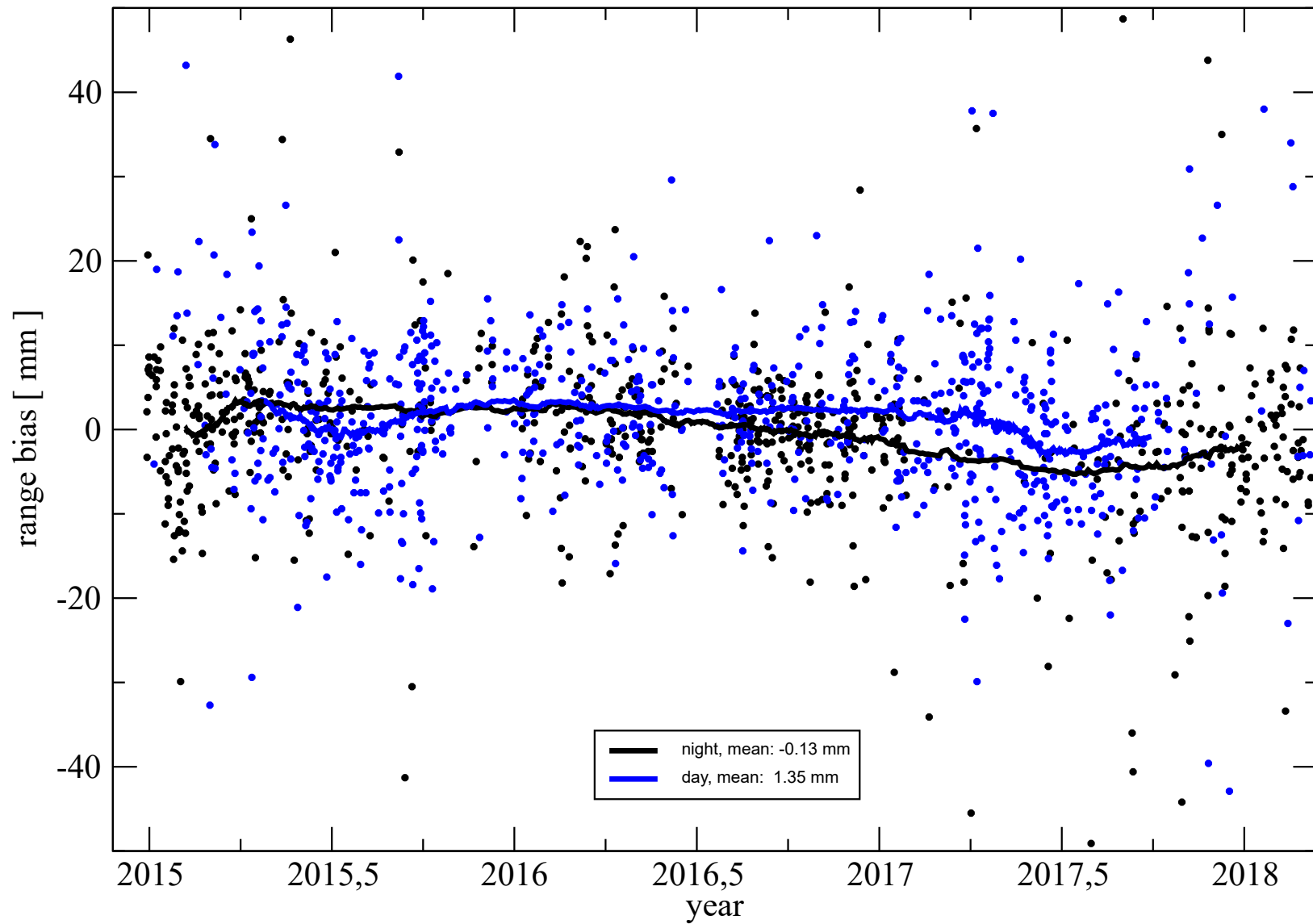
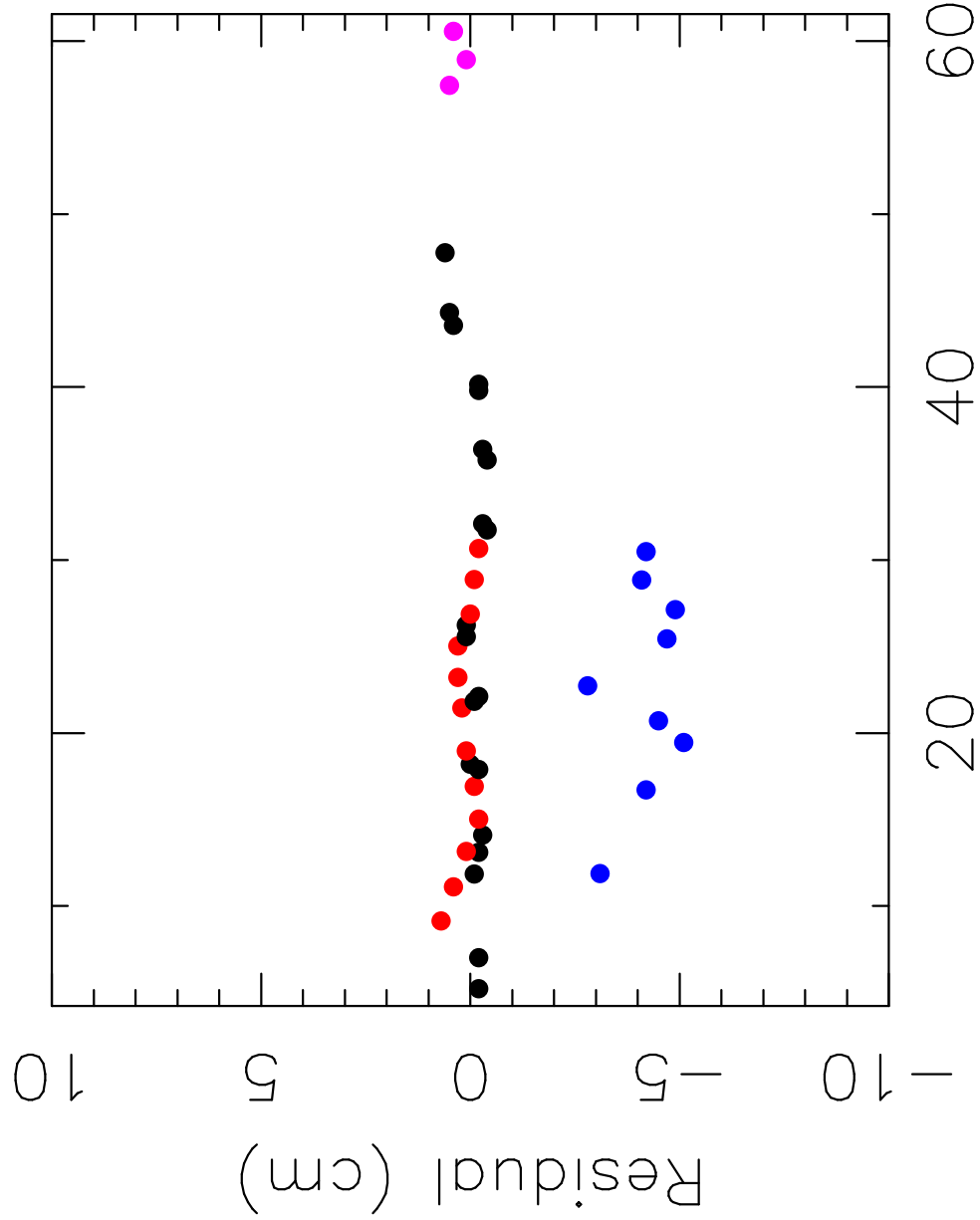


Fig. 8

2018-Mar-21



Satellite Lageos-1

7840 Herstmonceux

7841 Potsdam

7811 Borowiec

7110 Monument Peak

Mins after 02 hours

Fig. 9

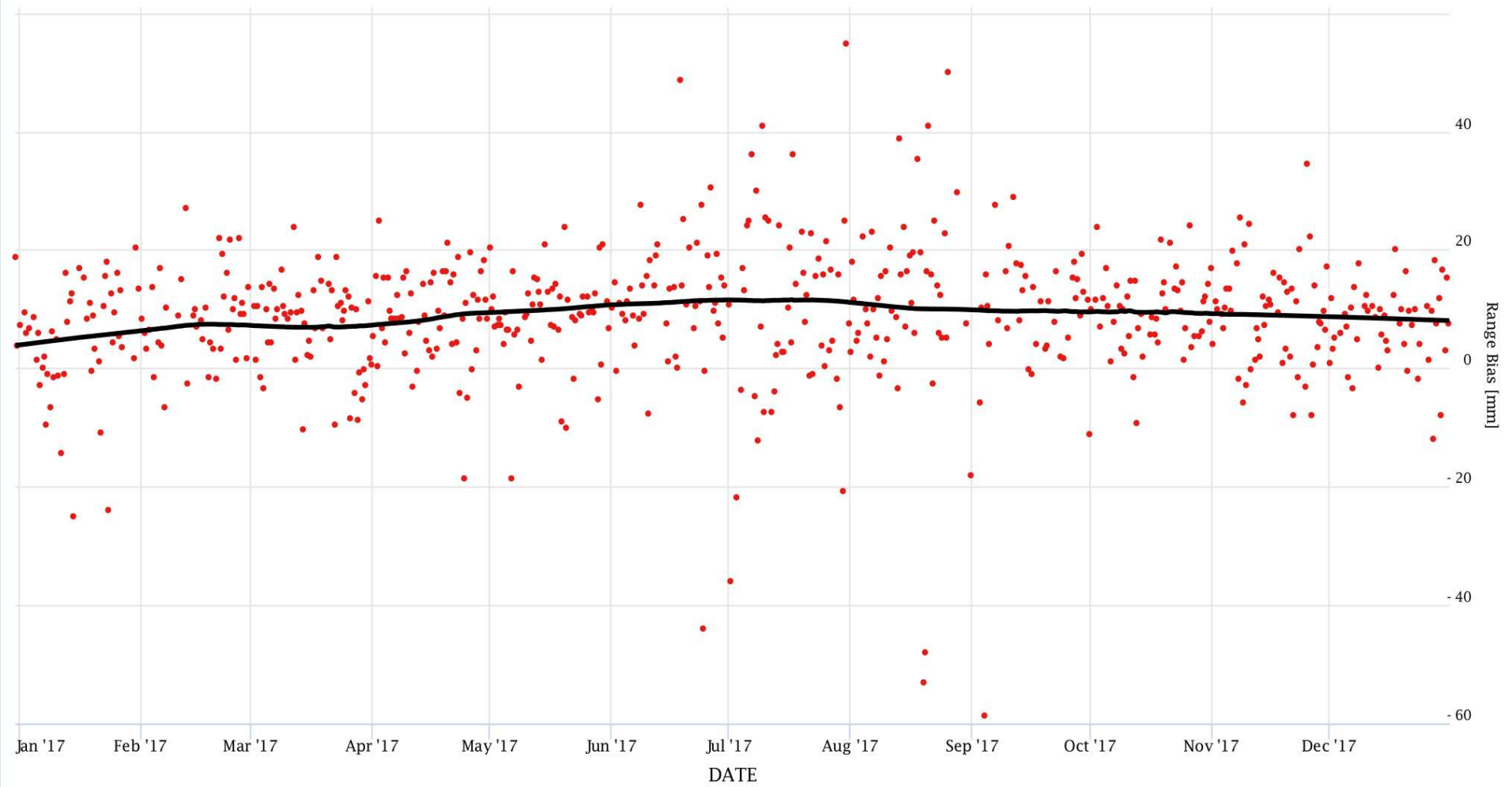
Range Bias Yarragadee 7090 AC(DGFI) LAGEOS2 SLRF2014



• Range Bias Yarragadee 7090 AC(DGFI) LAGEOS2
— LOESS Function 25 %

Zoom 1w 1m 6m **1y** 2y 3y All

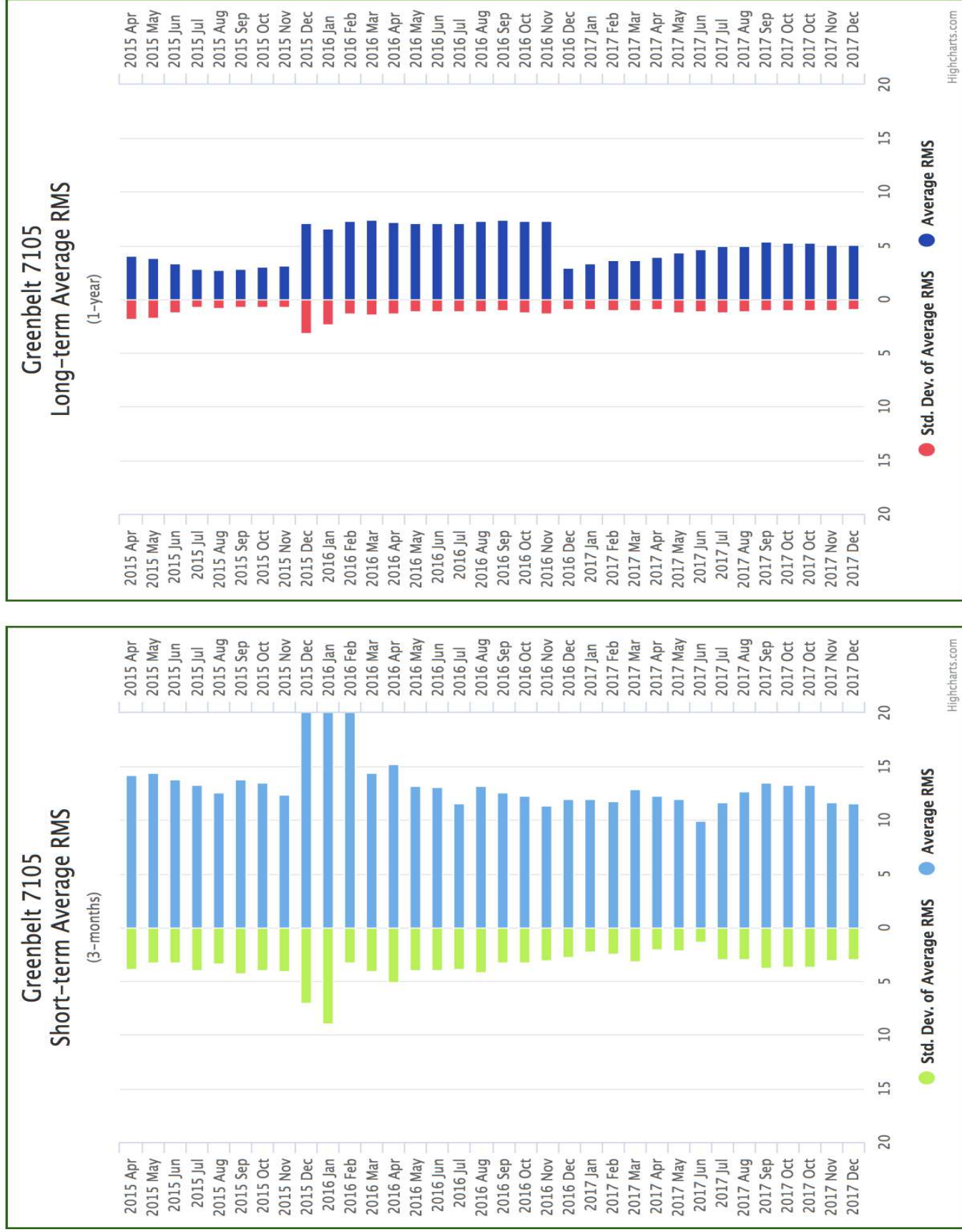
From Dec 31, 2016 To Dec 31, 2017



Highcharts.com

7090 Range Bias Yarragadee 7090 AC(DGFI) LAGEOS2
Mean/Std. Dev.: 9.01 ± 16.11 Count: 2,023

Greenbelt 7105



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Average RMS: It is computed from the input QC RMS's from the individual ACs that contribute to these series
 Std. Dev. of Average RMS: The statistical standard deviation of the "above" average RMS.